## falcON

#### a Cartesian FMM for the low-accuracy regime

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## **N**-body simulations in astronomy



HCG87: a group of galaxies



 $\omega {\rm Cen}:$  a globular cluster

## properties of stellar systems

- simple physics: Newtonian gravity
- very inhomogeneous
  - ⇒ large dynamic range
- $\triangleright$  dynamically young ( $t_{dyn} \simeq Myr-Gyr$ )
- well approximated as ensembles of point masses
  - ⇒ well described as Hamiltonian systems
    - $(\Rightarrow$  need symplectic time integration)

$$\begin{split} \mathsf{H} &= \sum_{i=1}^{N} \frac{m_i}{2} \left[ \boldsymbol{v}_i^2 - \sum_{j \neq i} \frac{G m_j}{|\boldsymbol{x}_i - \boldsymbol{x}_j|} \right], \qquad \boldsymbol{v}_i = \dot{\boldsymbol{x}}_i = \frac{\boldsymbol{p}_i}{m_i} \end{split}$$
with  $N \simeq 10^{5-20}$ 

equation of motion in continuum (mean-field) limit:

$$0 = \frac{\mathrm{d}f}{\mathrm{d}t} = \frac{\partial f}{\partial t} + \frac{\partial f}{\partial x} \cdot v - \frac{\partial f}{\partial v} \cdot \frac{\partial \Phi}{\partial x}$$

#### collisionless Boltzmann equation (CBE)

ightarrow f(x, v, t): distribution function (density in phase space)

 $ightarrow \Phi(x)$  : mean-field gravitational potential

both are related via the **Poisson equation**:

$$\nabla^2 \Phi(\boldsymbol{x}) = 4\pi G \int d^3 \boldsymbol{v} f(\boldsymbol{x}, \boldsymbol{v}, t)$$

## two-body relaxation

How good is the continuum description?

- stellar encounters deflect trajectories
  - ⇒ stellar orbits get randomized
  - ⇒ Maxwellian velocity distribution
- two-body relaxation time:

 $t_{\rm relax} \simeq 0.1 \frac{N}{\log N} t_{\rm dyn}$ 

#### 1 collision-dominated stellar dynamics

- $ightarrow t_{
  m relax} \lesssim$  age of system
- $\Rightarrow$  continuum limit not applicable
- $\Rightarrow$  must simulate Hamiltonian directly:
  - ▷ force computation is  $\mathcal{O}(N^2)$ 
    - $\Rightarrow$  computational effort limits  $N \leq 10^5$
  - close encounters are important
    - $\Rightarrow$  time integration becomes tedious

#### 2 collisionless stellar dynamics

- $\triangleright t_{relax} \gg age of system$
- ⇒ continuum limit applicable
- ⇒ solve CBE & Poisson equation

## 'collisionless' N-body simulations

#### How to solve the CBE?



- $\triangleright$  *f* is 6D & very inhomogeneous
  - $\Rightarrow$  (Eulerian) grid methods are useless
  - ⇒ Lagrangian method ('method of characteristics'):
- ▷ sample N trajectories { $\mu_i, x_i, v_i$ } from f(x, v, t = 0)
- $\triangleright$  solve equations of motion  $\ddot{x}_i = -\nabla \Phi(x_i, t)$
- $\triangleright$  CBE:  $\mu_i = \text{const}$  along trajectories
  - $\Rightarrow f(x, v, t)$  is represented by  $\{\mu_i, x_i(t), v_i(t)\}$
  - $\Rightarrow$  *f* is unknown
  - $\Rightarrow$  moments of f can be estimated
  - $\Rightarrow N \ll N$  is numerical parameter
  - ⇒ artificial two-body relaxation

#### How to solve the Poisson equation?

$$\nabla^2 \Phi(\mathbf{x}) = 4\pi G \int d^3 v f(\mathbf{x}, \mathbf{v}, t)$$

- 1 grid techniques (FFT, multigrid):
  - ▷ fast:  $\mathcal{O}(n_{\text{grid}} \log n_{\text{grid}})$
  - $\triangleright$  periodic ( $\Rightarrow$  cosmology)
  - problem: inhomogeneity (but: adaptive multigrid)
- 2 basic functions (using  $Y_{lm}$ ):
  - ▷ fast:  $\mathcal{O}(N n_{\text{basis}})$
  - problems: central singularity, spherical symmetry
- 3 Greens-function approach:

$$\Phi(\boldsymbol{x},t) = -G \int d^{3}\boldsymbol{x}' d^{3}\boldsymbol{v} \frac{f(\boldsymbol{x}',\boldsymbol{v},t)}{|\boldsymbol{x}-\boldsymbol{x}'|}$$

- general & adaptive
- problem: *f* is unkown
- $\Rightarrow$  estimate ( $\epsilon$ : softening length)

$$\Phi(\boldsymbol{x}_i,t) pprox - \sum_{i \neq j} rac{G \, \mu_j}{\sqrt{[\boldsymbol{x}_i - \boldsymbol{x}_j(t)]^2 + \epsilon^2}}$$

#### force softening to

- $\triangleright$  optimize force estimate (since *f* is unknown)
- suppress (unphysically) close encounters
- ⇒ force-estimation error (unavoidable)



true gravity of Hernquist model



estimation error with  $N = 10^6$ 

## computing the forces

 $\triangleright$  Greens-function approach  $\rightarrow$  Hamiltonian:

$$\mathsf{H} = \sum_{i=1}^{\mathsf{N}} \frac{\mu_i}{2} \left[ v_i^2 - \sum_{j \neq i} \frac{G \,\mu_j}{\sqrt{|x_i - x_j|^2 + \epsilon^2}} \right]$$

- $\triangleright$  how to evaluate  $\Phi \& \nabla \Phi$ ?
- can tolerate approximation error << estimation error</li>
   use approximative methods
- 1 direct summation (not approximative):
  - ▷ slow:  $\mathcal{O}(N^2)$  (but: GRAPE)
  - ▷ (unnecessarily) accurate
  - used in collisional N-body codes
- 2 Barnes & Hut (1986) tree code:
  - $\triangleright$  use hierarchical tree (usually: oct-tree)  $\Rightarrow$  fully adaptive
  - $\triangleright$  fast(er):  $\mathcal{O}(N \log N)$
  - most common method in astrophysics
  - violates Newton's 3rd law
    - ⇒ total momentun not conserved
- 3 traditional fast multipole method (FMM):
  - $\triangleright$  use hierarchy of cartesian grids  $\Rightarrow$  not fully adaptive
  - $\triangleright$  compute gravity via spherical multipoles & complex  $Y_{lm}$

 $\Rightarrow$  numerics complicated & cumbersome

formally O(N), but slower than tree code (for astrophyiscal applications, see Capuzzo-Colcetta & Miochi, 1998, JCP, 143, 29)



approximation error with  $N = 10^6$ 

## details of the tree code

#### 1 preparation phase

## 1.1 build a hierarchical tree of cubic cells ▷ cost: O(N log N)

- 1.2 pre-compute multipole moments etc
- 2 force computation: 'tree-walk'
  - ▷ for each body: compute force due to root cell
  - ▷ to compute force from cell:
    - if body is **well-separated** from cell:

compute force from multipole moments

otherwise

sum forces from daughter cells (recursive)

▷ cost:  $\mathcal{O}(\log N)$  per body  $\Rightarrow \mathcal{O}(N \log N)$ 

▷ the tree code is wasteful:

forces of neighbours are similar yet independently computed

## details of the FMM

here I describe traditional Greengard & Rokhlin (1987) FMM

#### 1 preparation phase

- 1.1 build a hierarchy of cartesian grids
- 1.2 pre-compute multipole moments etc (upward pass)
- 2 force computation

#### 2.1 interactions

- on each grid level:
- perform 'intermediate-field' interactions: compute & accumulate multipoles of gravity field

#### 2.2 downward pass

- ▷ pass field-multipoles down the hierarchy
- compute forces on finest grid

theoretical O(N) not demonstrated in practice
 not competetive with tree code in low-accuracy regime

## details of falcON

- ▷ hybrid of tree code & FMM
- takes the better of each method
- 1 preparation phase (as for tree code)
- 1.1 build a hierarchical tree of cubic cells
   ▷ cost: O(N log N)
- 1.2 pre-compute multipole moments etc
- 2 force computation

#### 2.1 interaction phase

- 'catch' all body-body interactions in well-separated node-node interactions:
  - if node-node interaction is executable execute it: accumulate field tensors
  - otherwise

split it & continue with child interations (recursive)

 $\triangleright$  cost: (better than)  $\mathcal{O}(N)$ , dominates

#### 2.2 evaluation phase

- pass field tensors down the tree
- compute forces at body positions
- $\triangleright$  cost:  $\mathcal{O}(N)$

 $ightarrow \sim 10$  times faster than tree code or FMM (at low accuracy)

### numerics of falcON

Wanted:

$$\Phi(\boldsymbol{x}_i) = -\sum_{j\neq i} \mu_j g(\boldsymbol{x}_i - \boldsymbol{y}_j),$$

Taylor expand g about  $\mathbf{R} = x_0 - y_0$ 

$$g(\boldsymbol{x}-\boldsymbol{y}) = \sum_{n=0}^{p} \frac{1}{n!} (\boldsymbol{x}-\boldsymbol{y}-\boldsymbol{R})^{(n)} \odot \boldsymbol{\nabla}^{(n)} g(\boldsymbol{R}) + \mathcal{R}_{p}(g),$$

Insert & sum over source cell B

$$\Phi_{\mathsf{B}\to\mathsf{A}}(x) = -\sum_{m=0}^{p} \frac{1}{m!} (x - x_0)^{(m)} \odot \mathsf{C}^{m,p} + \mathcal{R}_p(\Phi_{\mathsf{B}\to\mathsf{A}})$$
$$\mathsf{C}^{m,p} = \sum_{n=0}^{p-m} \frac{(-1)^n}{n!} \nabla^{(n+m)} g(R) \odot \mathsf{M}_{\mathsf{B}}^n,$$

$$\mathbf{M}_{\mathsf{B}}^{n} = \sum_{\boldsymbol{y}_{i} \in \mathsf{B}} \mu_{i} (\boldsymbol{y}_{i} - \boldsymbol{y}_{0})^{(n)}.$$

(Warren & Salmon 1995: Comp. Phys. Comm, 87, 266)

- $\sum_{m}$ : evaluation of gravity, represented by the **field tensors**  $\mathbb{C}^{m,p}$ , at position  $\boldsymbol{x}$
- $\sum_{n}$ : interaction between source cell B, represented by the **multipoles**  $M_{B}^{n}$ , and the sink cell A.

Difference to tree code:

- $\triangleright$  expansion in x (tree code:  $x \equiv x_0$ )
- mutuality of interactions

## gravity between well-seperated nodes



two well-separated cells

If  $|\mathbf{R}| > r_{A,crit} + r_{B,crit}$  with  $r_{crit} = r_{max}/\theta$ ,

 $\Rightarrow |x-y-R| < \theta |R| \forall x \in A, y \in B \&$  Taylor series converges force error of individual interaction:

$$\begin{aligned} |\boldsymbol{\nabla}\mathcal{R}_{p}(\boldsymbol{\Phi}_{\mathsf{B}\to\mathsf{A}})| &\leq \frac{(p+1)\theta^{p}}{(1-\theta)^{2}} \frac{\mathsf{M}_{\mathsf{B}}}{R^{2}} \\ &\propto \frac{\theta^{p+2}}{(1-\theta)^{2}} r_{\mathsf{B},\mathsf{max}}^{d-2} \propto \frac{\theta^{p+2}}{(1-\theta)^{2}} \,\mathsf{M}_{\mathsf{B}}^{(d-2)/d} \end{aligned}$$

> standard tree-code & FMM practice:  $\theta = \text{const}$ 

- ⇒ relative error controlled
- $\Rightarrow$  absolute error increases with M<sub>B</sub>
- ⇒ total error dominated by few interactions with large cells
- $\Rightarrow$  better:
  - $\triangleright$  balance **absolute** individual errors by  $\theta = \theta(M)$  with

$$\frac{\theta^{p+2}}{(1-\theta)^2} = \frac{\theta_{\min}^{p+2}}{(1-\theta_{\min})^2} \left(\frac{\mathsf{M}}{\mathsf{M}_{tot}}\right)^{(2-d)/d}$$

 $\Rightarrow$  reduce total error

#### accuracy vs. CPU time



mean (dashed) and 99 percentile (solid) relative force error

 $\varepsilon \equiv |a_{\text{approx}} - a_{\text{exact}}|/a_{\text{exact}},$ 

versus the CPU time (Pentium III/933Mhz in **2001**) for a galaxy (*left*) and a group f galaxies (*right*), sampled with (total)  $N = 10^4$  (*top*),  $N = 10^5$  (*mid-dle*), or  $N = 10^6$  (*bottom*). We used either  $\theta = \text{const}$  (*open triangles*) or  $\theta = \theta(M)$  (*solid squares*). The symbols along each curve correspond, from left to right, to values for  $\theta$  or  $\theta_{\min}$  of 0.2, 0.3, 0.4, 0.5, 0.6, 0.7, and 0.8.

### performance



CPU time per body (Pentium III/500Mhz in **2000**) versus N for a galaxy group.

#### what complexity?

▷ 8-folding  $N \Rightarrow N_I \rightarrow 8N_I + N_+$  and thus:

$$\frac{\mathrm{d}N_I}{\mathrm{d}N} \simeq \frac{N_I}{\mathrm{N}} \frac{\Delta \ln N_I}{\Delta \ln N} \approx \frac{N_I}{\mathrm{N}} + \frac{N_+}{N8 \ln 8},$$

with solution

$$N_I = c_0 N + \frac{N}{8 \ln 8} \int \frac{N_+}{N^2} \, \mathrm{d}N$$

▷ B&H tree code:  $N_+ \propto N$   $\Rightarrow N_I \propto N \log N$ ▷ Here:  $N_+(N)$  grows sub-linear at large N  $\Rightarrow N_I \propto N$ 

# comparison with other methods used in astrophysics



CPU time per body (**2001**) versus N for various techniques. Note that there are differences in the hard- & software, stellar system, and accuracy requirements.

by **2003/2004**: falcON is  $\sim$  3 times faster, but GRAPE-5 tree not.

## comparison with FMM

comparing under same conditions (bodies uniform in a cube)

$$E = \left[ \sum_{i} \left( \Phi_{i,\text{direct}} - \Phi_{i,\text{approx}} \right)^2 / \sum_{i} \Phi_{i,\text{direct}}^2 \right]^{1/2}$$

 $\triangleright$  low-accuracy regime:  $\sim 10$  times faster:

timing results (in seconds):

Ν	$T^a_{FMM}$	$T^a_{\rm direct}$	$E^a$	$T^b_{\mathrm{falcON}}$	$T^c_{\rm direct}$	$E^b$
20000	13.3	233	$7.9 imes10^{-4}$	0.97	136	$3.7 imes10^{-4}$
50000	27.7	1483	$5.2 imes10^{-4}$	2.64	924	$3.3 imes10^{-4}$
200000	158	24330	$8.4 imes10^{-4}$	10.77	14694	$3.4 imes10^{-4}$
500000	268	138380	$7.0 imes10^{-4}$	29.42	91134	$3.7 imes10^{-4}$
1000000	655	563900	$7.1 imes10^{-4}$	58.34	366218	$3.5 imes10^{-4}$

<sup>a</sup> FMM; data from Table I of Cheng et al. (1999: JCP, 155, 468)

<sup>b</sup> falcON on a computer identical to that used by Cheng et al.

<sup>c</sup> our own implementation of direct summation on the same computer

▷ high-accuracy regime:

falcON cannot compete with FMM

- $\Rightarrow$  accuracy & performance depend on both  $p \& \theta$ 
  - $\triangleright$  FMM: fixed ' $\theta$ ', vary p

 $\triangleright$  falcON: fixed p = 3, vary  $\theta$ 

 $\Rightarrow$  high accuracy requires higher order p

#### summary

▷ falcON = hybrid of tree code & FMM

▷ new features:

explicitly exploits mutuality of gravity

- ⇒ reduces computational effort
- ⇒ requires novel tree-walking algorithm
- ⇒ conservation of Newton's 3rd law
- mass-dependent  $\theta$ 
  - $\Rightarrow$  error balancing
  - $\Rightarrow$  reduces cost to **better** than  $\mathcal{O}(N)$
- $ightarrow \sim 10$  times faster than tree code or FMM

▷ publicly available

#### more dogmas

- balance errors
  - ⇒ reduce effort at given accuracy
- keep algorithm as simple as possible &
  - as complicated as necessary
  - $\Rightarrow$  high-order may be unnecessary
- write efficient code
  - $\Rightarrow$  avoid cache misses
    - $\Rightarrow$  data structure design
  - $\Rightarrow$  write generic code
  - $\Rightarrow$  do not rely too much on compliler optimization
    - ⇒ template metaprogramming